

## 2 Bounds on the viscosity coefficient of continental lithosphere from

- <sup>3</sup> removal of mantle lithosphere beneath the Altiplano and
- 4 Eastern Cordillera

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[1] The rapid rise of the central Andean plateau 8 9 between  $\sim 10$  and 6.8 Ma implies that mantle lithosphere, including eclogitized lower crust, was 10 removed from beneath the region in that time interval; 11 we infer from that removal that the average viscosity 12coefficient of mantle lithosphere was quite low when 13removal occurred. Using scaling laws for the growth 14of perturbations to the thickness of a dense layer over 15an inviscid substratum (Rayleigh-Taylor instability), 16we place bounds on the average viscosity coefficient 17for central Andean lithosphere. When compared with 18 laboratory measurements of flow laws for olivine and 19 eclogite, the allowed range of viscosity coefficients 20vields bounds on the temperature of  $\sim 500-800^{\circ}$ C at 21the Moho beneath this region and suggests that mean 22stresses across mantle lithosphere during continental 23deformation are less than  $\sim 50$  MPa. This range of 24temperature is comparable with, if a slightly lower, 25than we might expect for lithosphere approximately 26doubled in thickness and not yet equilibrated with the 27doubled crustal radioactivity. The mean deviatoric 28stress is comparable to that associated with stresses 29 that drive plates and hence shows that lithospheric 30 material is not too strong to prevent removal of its 31mantle part. Citation: Molnar, P., and C. N. Garzione (2007), 32 Bounds on the viscosity coefficient of continental lithosphere 33from removal of mantle lithosphere beneath the Altiplano and 34 Eastern Cordillera, Tectonics, 26, XXXXXX, doi:10.1029/ 352006TC001964. 36

### 38 1. Introduction

Removal of part or all of the mantle lithosphere offers a mechanism that can account for rapid changes in elevation of the overlying terrain and for rapid warming of the crust and remaining mantle lithosphere. To reconcile field observations of different kinds, many have appealed to some form of this process either delamination of crust from

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mantle lithosphere [Bird, 1978, 1979] or convective insta- 45 bility [e.g., England and Houseman, 1989; Houseman et al., 46 1981]. Although evidence consistent with such hypothe- 47 sized processes exists, convincing demonstrations that such 48 removal of mantle lithosphere has occurred have been 49 harder to find. Many doubt that removal can occur, because 50 the high viscosity of cold mantle lithosphere is thought to 51 prevent it [e.g., Morency et al., 2002; Schott and Schmeling, 52 1998]. Accordingly, a demonstration that this process has 53 occurred requires evidence that cannot be explained by 54 another process. We contend that the Altiplano of the central 55 Andes offers the most convincing evidence for removal of 56 mantle lithosphere, and we show that the implied average 57 viscosity of the mantle lithosphere beneath the Altiplano, 58 before it was removed, is consistent with laboratory experi- 59 ments on the flow of olivine or eclogite at temperatures that 60 are reasonable for the Altiplano at the time when dense 61 material was removed. 62

[3] We carry out an analysis similar to that of Molnar and 63 Jones [2004], who discussed the average viscosity beneath 64 the Sierra Nevada before mantle lithosphere was removed 65 from that area between  $\sim 10$  and 3.5 Ma. A variety of 66 observations suggest that eclogite-rich mantle lithosphere 67 was present beneath the Sierra at 10-12 Ma, but absent at 68 3.5 Ma. The most compelling evidence comes from a 69 change in xenolith composition from eclogite and garnet 70 pyroxenite prior to  $\sim 8$  Ma to a spinel peridotite composition 71 since ~1 Ma [Ducea and Saleeby, 1996, 1998; Lee et al., 72 2000, 2001]. These xenoliths, derived from 40 to 100 km 73 depths, suggest that eclogitic lower crust and mantle litho-74 sphere were removed. The eruption of potassic basalts at 75 3.5 Ma is also consistent with removal of mantle lithosphere 76 and provides a minimum age of  $\sim$ 3.5 Ma on the timing of 77 removal [Farmer et al., 2002]. 78

# **2. Removal of Eclogite-Rich Mantle**79Lithosphere From Beneath the Altiplano80

[4] The central Andean plateau (Figure 1), with a width 81 of  $\sim$ 400 km and an average elevation of  $\sim$ 4 km, is the 82 second largest high plateau on Earth after the Tibetan 83 plateau. In the central Andes where the plateau is widest, 84 the Western and Eastern Cordillera, with peak elevations 85 exceeding 6 km, bound the internally drained Altiplano 86 basin whose average elevation is  $\sim$ 3800 m. Active magma-87 tism characterizes the Western Cordillera. The Eastern 88 Cordillera and Altiplano basin preserve a history of folding 89 and faulting. Constraints on the timing and distribution of 90

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**Figure 1.** Elevations of the central Andean plateau between  $12^{\circ}$ S and  $27^{\circ}$ S, constructed with SRTM30 data set. Regions outlined in red show the extent of middle-late Miocene low-relief paleosurfaces that underwent rotation and incision beginning in late Miocene time. Green dots show locations of Miocene paleobotanical estimates of paleoelevation. Blue dots show locations of late Miocene paleoelevation estimates from  $\delta^{18}$ O and  $\Delta_{47}$  of authigenic carbonates shown in Figure 2.

deformation indicate that the Andean plateau, from the 91 monoclinal structure that forms the western flank through 92the Altiplano and Eastern Cordillera, underwent horizontal 93shortening between  $\sim 40$  and  $\sim 7$  Ma [Barnes et al., 2006; 94Carrapa et al., 2005, 2006; Deeken et al., 2006; Ege et al., 952007; Elger et al., 2005; Farías et al., 2005; Gillis et al., 96 2006; Horton, 2005; Horton and DeCelles, 1997; Klev, 97 1996; Kraemer et al., 1999; McQuarrie, 2002; Müller et al., 982002; Sheffels, 1990; Victor et al., 2004]. 99

100 [5] In the context of the shortening history of the central 101 Andes, several authors have used geomorphic observations 102 to infer that the surface of this area rose significantly since 103  $\sim$ 10 Ma [e.g., *Gubbels et al.*, 1993; *Isacks*, 1988; *Kennan et 104 al.*, 1997]. Widespread, low-relief paleosurfaces on both the 105 eastern slope of Eastern Cordillera [*Gubbels et al.*, 1993; 106 *Kennan et al.*, 1997] and the western slope of the Western Cordillera [Farías et al., 2005; García and Hérail, 2005; 107 Hoke et al., 2004; Kober et al., 2006; Schlunegger et al., 108 2006; von Rotz et al., 2005; T. F. Schildgen et al., Tectonics 109 of the western margin of the Altiplano in Southern Peru 110 from river incision history, unpublished manuscript, 2006; 111 G. D. Hoke et al., Geomorphic evidence for post-10 Ma 112 uplift of the western flank of the central Andes  $(18^{\circ}30' - 113)$ 22°S), submitted to Tectonics, 2007, hereinafter referred to 114 as Hoke et al., submitted manuscript, 2007] reflect the 115 remnants of low-relief drainage systems between  $\sim$ 7 Ma 116 to 12 Ma in the Eastern Cordillera and until  $\sim 10$  Ma on the 117 western slope, after which widespread incision of both the 118 eastern and western paleosurfaces had begun, in some areas 119 more tightly dated to be before  $\sim 6.5$  Ma. Reconstructions 120 of the relief in these drainage systems have been used to 121 infer  $\sim 1$  to 2 km of surface uplift of the flanks of the 122



Figure 2. History of elevation change from multiple proxies in the northern Altiplano and Eastern Cordillera. The location of each record is shown on Figure 1. Paleobotany estimates from Gregory-Wodzicki et al. [1998] and Gregory-Wodzicki [2000]; Oxygen isotope estimates from Garzione et al. [2006]; and  $\Delta_{47}$  estimates from Ghosh et al. [2006].

Andean plateau [e.g., Kennan et al., 1997; Hoke, 2006; 123124Kober et al., 2006; Hoke et al., submitted manuscript, 1252007]. Finally, concurrent with the rise of the Andean plateau, the locus of active crustal shortening migrated to 126the sub-Andes on the eastern margin of the belt after 127~10 Ma [e.g., Echavarria et al., 2003; Moretti et al., 128 1996; Uba et al., 2006]. 129

[6] Quantitative estimates of paleoelevation come from 130marine deposits in the Altiplano basin, fossil leaf physiog-131 nomy, oxygen isotopes from carbonate sediment, and  $\Delta_{47}$ 132paleothermometry (Figure 2). Shallow marine deposits of 133 the 70-60 Ma El Molino Formation require that the 134Altiplano lay at sea level at the end of Cretaceous time 135[Sempere et al., 1997]. Paleotemperature estimates derived 136137 from fossil leaf physiognomy in the northern Altiplano and 138 Eastern Cordillera (Figure 1) suggest paleoelevations of no more than a third of the plateau's modern average height of 139 $\sim$ 4 km at  $\sim$ 15 to 20 Ma [Gregory-Wodzicki, 2000] and no 140 more than half by  $\sim 10$  Ma [Gregory-Wodzicki et al., 1998]. 141 Recent oxygen-isotope paleoaltimetry [Garzione et al., 1422006] and  $\Delta_{47}$  paleothermometry [Ghosh et al., 2006] 143 suggest  $3 \pm 1$  km between  $\sim 10$  Ma and 6.8. Ma. 144 [7] Rapid surface uplift of several kilometers of a region 145

as wide as the Andean plateau in  $\sim 3 \pm 1$  Myr of the late 146 Miocene Epoch (Figure 2) reflects a rate of surface uplift 147 that is too high to be generated by crustal shortening alone 148149and requires the removal of relatively dense eclogite and mantle lithosphere [Garzione et al., 2006]. Flow of middle-150lower crust [e.g., Gerbault et al., 2005; Hindle et al., 2005; 151152Husson and Sempere, 2003] may have redistributed crustal material and contributed to the nearly flat surface of the 153154Altiplano. Lower crustal flow, however, cannot explain the 155simultaneous rise not only of the Altiplano but also of both the Eastern and Western Cordilleras some 300 km apart, for 156

there is no obvious source of the volume of crustal material 157 that must be injected into crust beneath these regions. In 158 fact, crustal thickening in the Eastern and Western Cordil- 159 leras associated with crustal shortening and magmatism 160 should result in a source of excess lower crustal material 161 from the cordilleras, not injection of material beneath them. 162

[8] Constraints on the thermal and structural character- 163 istics of the crust and mantle as well as the volcanic history 164 are consistent with removal of mantle lithosphere and 165 eclogitic lower crust from below the Andean plateau. The 166 crustal thickness in the central Andes varies between 70 km 167 below the highest topography in the Eastern and Western 168 Cordilleras and 59 to 64 km in the central Altiplano, 169 suggesting that the region is in approximate Airy isostatic 170 balance [Beck and Zandt, 2002; James, 1971; Yuan et al., 171 2002]. Seismic tomography of the mantle between 16° and 172 20°S, however, shows the lowest P wave speeds below the 173 Altiplano-Eastern Cordillera transition and suggests that 174 virtually all of the mantle lithosphere has been removed in 175 this region [Dorbath and Granet, 1996; Myers et al., 1998]. 176 In addition, the crustal column beneath the Altiplano lacks 177 typical high-speed lower crust, suggesting a felsic compo- 178 sition that is typical of upper crust [Beck and Zandt, 2002]. 179 All of these observations support previous suggestions [Kav 180 and Mahlburg-Kay, 1991; Kay and Kay, 1993] that both 181 mantle lithosphere and eclogitic lower crust were removed 182 from below the Altiplano and the western part of the Eastern 183 Cordillera [Beck and Zandt, 2002]. The eruption of mafic 184 lavas throughout the northern and central Altiplano begin- 185 ning at  $\sim$ 7.5 to 5.5 Ma [Carlier et al., 2005; Lamb and 186 Hoke, 1997] and at  $\sim$ 7 to 3 Ma in the southern Altiplano 187 and Puna [Kay et al., 1994, 1999] have been inferred to 188 reflect Late Miocene to Pliocene removal of eclogitic lower 189 crust and mantle lithosphere beneath the Altiplano and Puna 190 plateaus [Kay et al., 1994; Lamb and Hoke, 1997]. 191 High <sup>3</sup>He/<sup>4</sup>He ratios across much of the Altiplano and 192 Eastern Cordillera, interpreted to result from degassing of 193 mantle-derived magmas, also support the inferred removal 194 of mantle lithosphere, in this case for virtually the entire 195 Altiplano [Hoke et al., 1994]. 196

#### 3. Rayleigh-Taylor Instability

overlies cold mantle lithosphere.

197 [9] The relatively low temperature of mantle lithosphere 198 makes this material in most regions denser than astheno- 199 sphere when at the same pressure, even where eclogite 200 comprises a negligible fraction of the lithosphere. The 201 notable exception is beneath Archean cratons, where chem- 202 ical differentiation has made mantle lithosphere intrinsically 203 less dense than asthenosphere at the same temperature [e.g., 204 Jordan, 1975; Poudjom Djomani et al., 2001]. Thus mantle 205 lithosphere in most regions should be unstable, and pertur- 206 bations to its thickness should grow with time, unless 207 diffusion of heat can erase the perturbations before they 208 can grow. The presence of eclogite enhances this instability 209 [e.g., Kay and Mahlburg-Kay, 1991; Nelson, 1991; Jull and 210 Kelemen, 2001], particularly where a thick layer of eclogite 211

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t1.1 Table 1. List of Symbols

t1.2	Symbol	Definition	
+1 9	A	empirical constant relating strain rate to stress difference in laboratory measurements of high temperature flow.	
61.0 +1.4	D	viscosity coefficient for non Newtonian viscosity	
11.4	D	viscosity coefficient for non-Newtonian viscosity	
	C	the growth rate of an instability for non-Newtonian viscosity; C depends on the wavelength of the	
t1.5		perturbation and weakly on <i>n</i>	
t1.6	d	thickness of cooling layer	
t1.7	Ε	second invariant of the strain rate tensor	
t1.8	g	gravity	
t1.9	$H_a$	activation enthalpy	
t1.10	L	thickness of unstable layer (lithosphere)	
t1.11	$\Delta L_0$	perturbation to the thickness of unstable layer (lithosphere	
t1 12	n	power that relates strain rate to deviatoric stress in the	
+1 13	R	universal as constant	
01.10	4	alansed time between initiation of a perturbation and	
t1.14	ι <sub>b</sub>	when perturbations sinks to infinite depth	
t1.15	Т	temperature (in kelvins)	
	$\Delta \rho$	density difference between unstable top layer (lithosphere)	
t1.16		and underlying layer (asthenosphere)	
t1.17	$e_{ii}$	strain rate tensor	
t1.18	<i>к</i>	coefficient of thermal diffusivity	
t1.19	$ au_{ij}$	deviatoric stress tensor	

[10] The nonlinear relationship between deviatoric stress 213and strain rate that applies to most rock-forming minerals 214can make lithosphere more stable than it would be if 215deformation of the lithosphere obeyed Newtonian viscosity 216[e.g., Conrad and Molnar, 1999; Houseman and Molnar, 2172001]. In the case of the Altiplano, however, where the crust 218has been thickened to roughly twice its normal value, this 219large perturbation to the thermal structure should be suffi-220cient to prevent stabilization in a period shorter than tens of 221millions of years. The time constant for decay of thermal 222perturbations by diffusion of heat in a layer of thickness d is 223given by  $d^2/\kappa$  where  $\kappa$  is the coefficient of thermal diffu-224sivity ( $\sim 10^{-6}$  m<sup>2</sup> s<sup>-1</sup>). For a layer 70 km thick (the 225maximum crustal thickness beneath the Andes, and much 226thinner than thickened lithosphere), the thermal time con-227stant is 160 Myr, much longer than the time interval over 228which material was removed. Thus we need not consider 229diffusion of heat, and we can use scaling laws based on 230231Rayleigh-Taylor instability, which results when an intrinsically dense layer overlies a less dense fluid. 232

[11] By exploiting the time for perturbations to grow by 233Rayleigh-Taylor instability as a mechanism for removal of 234lithosphere, we assume that the material that descends, first, 235does so in narrow plumes or sheets, and second, draws 236material from neighboring lithosphere. Thus we treat the 237understanding and the timescale for the downwelling of 238material as surrogates for those of removal of material from 239adjacent regions. As mass must be conserved, if some 240lithosphere descends, other, surely hotter, and therefore less 241dense material must rise to replace it. Because many factors 242can affect the distance scales involved in such removal, and 243much understanding of Rayleigh-Taylor instability remains 244

to be gained, we do not try to specify where downwelling 245 occurred or what its planform might have been. 246

#### 3.1. Scaling Laws

[12] Using different approaches, Canright and Morris 248 [1993] and Houseman and Molnar [1997] obtained scaling 249 laws for the time that must elapse for a perturbation in the 250 thickness of an unstable layer to sink to infinite depth. 251 Canright and Morris [1993] treated the layer as a thin 252 viscous sheet, and hence with no shear stress on vertical or 253 horizontal planes, including at the top and bottom of the 254 sheet. Houseman and Molnar [1997] considered the com- 255 plete deformation field of a layer with a rigid top boundary. 256 The extent to which mantle lithosphere (with or without 257 eclogite) is detached from the overlying crust or tightly 258 attached to it remains controversial, but by considering 259 these extremes, we span the range of likely top boundary 260 conditions on the mantle lithosphere. Although the analysis 261 by Canright and Morris [1993] does not lend itself to 262 consideration of additional complexity in the distribution 263 of density or viscosity, numerical experiments by others 264 [e.g., Jull and Kelemen, 2001; Molnar et al., 1998] allow 265 assessments of some such complexity. Also, Conrad and 266 Molnar [1999] showed that convective instability, not only 267 with vertically varying density and viscosity, but also with 268 diffusion of heat, obeys a scaling law similar to that found 269 by Houseman and Molnar [1997] for Rayleigh-Taylor 270 instability. Thus ignoring diffusion of heat and considering 271 Rayleigh-Taylor instability should introduce a negligible 272 error to our estimates of viscosity coefficient. 273

[13] Most treatments of convection within the mantle 274 beneath the lithosphere consider Newtonian viscosity, but 275 in the lithosphere, where temperatures are relatively low, we 276 must allow for a non-Newtonian constitutive relationship 277 between stress and strain rate. As discussed below, labora- 278 tory measurements of rock-forming minerals suggest a 279 relationship between deviatoric stress,  $\tau_{ij}$ , and strain rate, 280  $e_{ij}$ , that can be written as 281

$$\tau_{ij} = BE^{(1-n)/n} e_{ij} \tag{1}$$

*B* is a temperature-dependent (and pressure-dependent) 283 viscosity coefficient (with units of Pa s<sup>1/n</sup>), and  $E^2 = (1/2)$  284  $\sum_{i,j} e_{ij} e_{ij}$  is the second invariant of the strain rate tensor,  $e_{ij}$ , 285 and for olivine and most earth-forming minerals, the exponent 287 n = 3-3.5. (Symbols are defined in Table 1.)

[14] The scaling laws developed by *Canright and Morris* [1993] and *Houseman and Molnar* [1997] consider constitutive laws like (1), but expressed with dimensionless quantities. To use these theoretical analyses and numerical experiments, we consider dimensionless distances obtained by dividing distances by lithospheric thickness, *L*, and dimensionless times, t', scaled using

$$t' = t(\Delta \rho g L/B)^n \tag{2}$$

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in which  $\Delta\rho$  is the difference between the density of the  $_{297}$  unstable layer (lithosphere) and that of the substratum (as- $_{298}$  thenosphere), and g is gravity.

[15] To approximate the behavior of an unstable layer, 299Canright and Morris [1993] used a dense thin viscous 300 sheet, in which shear stresses on horizontal and vertical 301 planes are negligible. This assumption treats any substratum 302 as inviscid and the top boundary as free of shear stress. 303 Moreover, to obtain an analytic solution, they assumed a 304constant  $\Delta \rho$  within the layer. They derived an expression 305that can be integrated to give the time,  $t'_b$ , that a perturbation 306 to the thickness of the sheet,  $\Delta L'_0 = \Delta L_0 / L$  (where  $\Delta L_0$  is 307 the amplitude of the dimensional perturbation), grows to 308 infinite thickness,  $\Delta L' \rightarrow \infty$  (or thins to zero thickness). For 309 n = 3, an approximate value of n for olivine, they obtained 310 an analytic solution for  $t'_{h}$ : 311

$$t'_{b}(\Delta L'_{0}) = \frac{1 + \Delta L'_{0}}{4\left[\left(1 + \Delta L'_{0}\right)^{2} - 1\right]^{2}} + \frac{1 + \Delta L'_{0}}{8\left[\left(1 + \Delta L'_{0}\right)^{2} - 1\right]} + \frac{1}{16}\ln\frac{\Delta L'_{0}}{2 + \Delta L'_{0}}$$
(3)

Not surprisingly, numerical integration of their governing equation for n = 3.5 yields a solution that does not differ much from that for n = 3 [*Molnar and Jones*, 2004].

<sup>316</sup> [16] Using a less elegant approach than did *Canright and* <sup>317</sup> *Morris* [1993], *Houseman and Molnar* [1997] considered a <sup>318</sup> dense sheet with no horizontal or vertical movement of its <sup>319</sup> top surface (rigid top boundary) and found that the elapsed <sup>320</sup> time for a dimensionless harmonic perturbation  $\Delta L'_0$  to <sup>321</sup> grow to infinite depth could be approximated by

$$t'_b(\Delta L') = \left(\frac{n}{C}\right)^n \frac{\Delta L'^{1-n}_0}{n-1} \tag{4}$$

Here *C* is an empirically determined constant that is  $\sim 0.76$ for n = 3 in a layer of constant density (and using the definition of *E* given above). *Houseman and Molnar* [1997] interpolated between estimates of *C* from numerical experiments for n = 3 and n = 5 to suggest that for n = 3.5,  $C \approx$ 0.9.

17] Although the different forms of (3) and (4) might suggest markedly different dependences of  $t'_b$  on  $\Delta L'_0$ , *Molnar and Jones* [2004] showed them to be quite similar at least for n = 3 and n = 3.5. Thus, if we can assign values not only to  $\Delta \rho$ , g, and L in (2), but also to  $t_b$  and  $\Delta L'_0$ , we can combine (3) or (4) with (2) to deduce a value of B from

$$B = \Delta \rho g L \left(\frac{t_b}{t'_b}\right)^{\frac{1}{n}} \tag{5}$$

#### 337 3.2. Application of Scaling Laws to the Andean Plateau

<sup>338</sup> [18] With a crustal thickness of 70 km for the Eastern <sup>339</sup> Cordillera, roughly twice that of typical crust, we may infer <sup>340</sup> that mantle lithosphere had also thickened approximately <sup>341</sup> two times; hence  $\Delta L'_0 \approx 1$ . If n = 3, (3) yields  $t'_b = 0.07$ , <sup>342</sup> and (4) gives  $t'_b = 30.8$ . With n = 3.5, appropriate for olivine <sup>343</sup> or eclogite, an integration of *Canright and Morris*'s expres-<sup>344</sup> sions yields  $t'_b = 0.1$ , and with (4)  $t'_b = 46.4$ . The large differences in values of  $t'_b$  result from the resistance to flow 345 imposed by the top boundary in the case considered by 346 *Houseman and Molnar* [1997] and the freedom of a layer to 347 slip horizontally if its top is stress-free. 348

[19] As discussed above, the Andean plateau rose within 349 an interval of  $\sim 3 \pm 1$  Myr. We treat that interval as defining 350 the duration between the time when a large perturbation to 351 the thickness of the lithosphere grew to effectively infinite 352 depth, and hence as defining  $t_b = 3 \pm 1$  Myr. Although this 353 process strictly applies to the time needed for a perturbation 354 in thickness to grow to great depth, because the material that 355 sinks must be drawn from the adjacent lithosphere, we treat 356 that interval of time as also defining the period in which 357 lithosphere is removed from beneath the Andean plateau. 358 Also as noted above, the recent rise of the Altiplano of  $3 \pm 1$  km, 359 if compensated by removal of mantle lithosphere gives an 360 estimate of the product  $\Delta \rho L = 8.4(\pm 2.8) \times 10^6$  kg m<sup>-2</sup>. 361 This combination of parameters suggests that for a rigid top 362 boundary  $B \approx 9.1 \times 10^{12}$  Pa s<sup>1/3</sup> (for n = 3) or  $B \approx 1.57 \times 363$ 10<sup>12</sup> Pa s<sup>1/3.5</sup> (n = 3.5) and for a free top  $B \approx 1.20 \times 10^{12}$  Pa s<sup>1/3</sup> (n = 3) and  $B \approx 2.7 \times 10^{11}$  Pa s<sup>1/3.5</sup> (n = 3.5). 364365

#### **3.3.** Discussion of Uncertainties

[20] Several sources contribute to uncertainty in B. First, 367 consider the effect of errors in assumed parameters (see 368 Table 2). Because the values of  $t'_b$  and  $t_b$  contribute to 369 estimates of B in (5) only when the cube or 3.5th root is 370taken, their uncertainties are not very important. For in- 371 stance, the 1-Myr uncertainty in the time for deblobbing 372 contributes less than a 10% error to the estimates of B. If we 373 assumed that the initial perturbation were 50% ( $\Delta L'_0 \approx 0.5$ ) 374 instead of 100%, the estimates of B for either free slip or 375 rigid top boundaries would be 65% smaller, than those 376 given above. Finally, the 1-km (33%) uncertainty in the 377 elevation change between 10 and 6.8 Ma implies a 33% 378 uncertainty in the product  $\Delta \rho L$  in (5), and a corresponding 379 uncertainty in B. Thus the combined uncertainties in the 380 assumed parameters leads to an uncertainty in *B* that is less 381 than 100% (less than a factor or 2) and closer to 50%. 382

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[21] Assumptions of boundary conditions and in the 383 applicability of scaling laws derived for idealized structures 384 cause the largest uncertainties in estimates of *B*. The values 385 of *B* for free slip at the top and a rigid top differ by a factor 386 of 7.6 or 5.8 depending on the value of *n*. Surely the top is 387 neither free nor rigid, and the likely value of *B* ought to 388 lie between these values:  $9.1 \times 10^{12}$  Pa s<sup>1/3.5</sup> and  $1.20 \times 389$   $10^{12}$  Pa s<sup>1/3.5</sup> for n = 3, or between  $1.57 \times 10^{12}$  Pa s<sup>1/3.5</sup> and 390  $2.7 \times 10^{12}$  Pa s<sup>1/3.5</sup> for n = 3.5.

[22] The scaling laws in (2) and (3) consider an inviscid 392 fluid beneath the unstable layer, but a finite viscosity of the 393 asthenosphere will retard growth of an unstable mantle 394 lithosphere. *Molnar et al.* [1998] ran numerical experiments 395 for cases in which the viscosity coefficient decreased 396 exponentially across the unstable layer to a constant value 397 equal to that of the lower layer and found that the presence 398 of a viscous substratum decreased values of *C* in (3) by 399 about 20% below those with an inviscid substratum. Thus 400 ignoring the viscosity of the substratum makes our estimates 401 of *B* too large by ~20%. Also we use the elapsed time for 402

t2.1 Table 2. Summary of Contributions to Uncertainties

t2.2	Assumption in Error	Effect on Estimated Value of B	
t2.3	$30\%$ error in $t_b$	uncertain by <10%	
t2.4	Smaller perturbation, $\Delta L'_0 \approx 0.5$ (not 1.0)	smaller by 65%	
t2.5	1-km (33%) error in the change in elevation	uncertain by $33\%$ overestimated by 5.8 to 7.6 times underestimated by 5.8 to 7.6 times overestimated by ~20%	
t2.6	No shear stress on top		
t2.7	No slip on top		
t2.8	Negligible asthenospheric viscosity		
t2.9	Descent to 300–400 km (not infinite depth)	underestimated by $\sim 5\%$	
t2.10	Constant density in lithosphere	overestimated by 20-30%	

403 the unstable layer to sink to infinite depth, but were we to 404 use a depth of 300-400 km, we would augment the 405 estimated values of *B* by only 5% [*Molnar and Jones*, 406 2004].

[23] The scaling laws also are based on theory or exper-407iment for which the density anomaly is constant throughout 408the layer, but if the density anomaly decreases from a 409maximum at the top of the layer to zero at its bottom, 410growth is slower. Houseman and Molnar's [1997] numer-411 ical experiments for such a density distribution yield values 412of  $C \sim 20-30\%$  smaller than those for constant density. 413Such a decrease in C leads to a comparably smaller inferred 414viscosity coefficient B. 415

[24] Finally, the assumption of a constant viscosity coef-416 ficient in the layer is unrealistic. A more realistic viscosity 417 coefficient that decreased with depth in the layer would 418make the bottom part sink rapidly, but the upper part more 419slowly than the average value. Correcting for such a 420decrease with depth is difficult. Houseman's finite element 421code, which solves the equations of motion in a Lagrangian 422 frame, becomes numerically unstable when elements de-423form from roughly equidimensional to elongated triangles, 424 as they do when the large strains develop as blobs sink to 425great depth and the remaining lithosphere thins. We cannot 426quantify easily the error introduced by using the scaling 427 laws to estimate the average viscosity of the lithosphere, but 428 the numerical experiments on convective instability by 429Conrad and Molnar [1999] show that the same scaling 430laws used here apply to those cases for which viscosity 431decreased exponentially with temperature and hence ap-432proximately exponentially with depth through the unstable 433434layer.

435 [25] The arguments given above suggest that allowance 436 for a linearly decreasing density anomaly across the layer 437 and of a finite viscosity of the substratum would make the 438 average value of *B* for the Altiplano ~50% smaller than the 439 values given above, and uncertain by ~50%. Thus, if n =440 3.5 we use  $B \approx 7.9 (\pm 3.9) \times 10^{11}$  Pa s<sup>1/3.5</sup> for a free top and 441  $B \approx 1.4 (\pm 0.7) \times 10^{11}$  Pa s<sup>1/3.5</sup> for a rigid top, or if n = 3, 442  $B \approx 4.6 (\pm 2.3) \times 10^{12}$  Pa s<sup>1/3</sup> for a free top boundary and  $B \approx$ 443  $6.0 (\pm 3.0) \times 10^{11}$  Pa s<sup>1/3</sup> for a rigid top.

#### 445 **4. Discussion**

446 [26] Because few readers have experience with the vis-447 cosity coefficient *B* (or its units of Pa s<sup>1/3</sup> or Pa s<sup>1/3.5</sup>), we 448 address the significance of the range of inferred values given above in two ways: what they imply (1) for magni- 449 tudes of stress in the lithosphere and (2) for temperatures at 450 depth. First, *B* relates magnitudes of stress to strain rates as 451 shown in equation (1), and second, using laboratory meas- 452 urements of *B* as a function of temperature, we may use the 453 estimates given above to infer temperatures of the upper 454 mantle at the Moho. 455

[27] Much of the crustal thickening in the Altiplano and 456 Eastern Cordillera occurred between 40 and  $\sim 10$  Ma [e.g., 457 Elger et al., 2005; McQuarrie, 2002]. Suppose that crust 458 doubled in thickness in 30 Myr, so that the average strain 459 rate was  $e = 1 \times 10^{-15} \text{ s}^{-1}$ . If thickening developed by pure 460 shear so that horizontal compression occurred at the same 461 (but negative) rate, then  $E = 1 \times 10^{-15} \text{ s}^{-1}$  also. 462 Corresponding average deviatoric stresses across the mantle 463 lithosphere, given by  $\tau = Be^{1/n}$ , are 7.3  $\leq \tau \leq 41$  MPa for 464 1.4  $\leq B \leq 7.9 \times 10^{11}$  Pa s<sup>1/3.5</sup> and 6.0  $\leq \tau \leq 46$  MPa for 465 0.60  $\leq B \leq 4.6 \times 10^{12}$  Pa s<sup>1/3</sup>. Because these are averages 466 for the entire mantle lithosphere, maximum values should 467 be larger. These average bounds span those inferred for 468 Tibet [e.g., England and Molnar, 1997; Flesch et al., 2001], 469 where lithosphere has been thinned. Most important, they 470 are not so large as to prohibit deformation of mantle 471 lithosphere, for such average stresses are comparable to 472 those that drive plates [e.g., Chapple and Tullis, 1977; 473 Forsyth and Uyeda, 1975; McKenzie, 1972]. 474

[28] To compare these inferred values of *B* with those 475 derived from laboratory measurements, we first must take 476 into account the temperature dependence of *B*. Experimen-477 tally determined power law relationships between stress and 478 strain rate, such as those for olivine [e.g., *Goetze*, 1978; 479 *Karato and Wu*, 1993] or eclogite [*Jin et al.*, 2001], can be 480 written as [e.g., *Molnar et al.*, 1998] 481

$$B(T) = 3^{-\frac{n+1}{2n}} \left(\frac{A}{2}\right)^{-\frac{1}{n}} \exp\left(\frac{H_a}{nRT}\right)$$
(6)

where *A* and *n* are parameters determined experimentally in 483 the laboratory,  $H_a$  is the activation enthalpy, *T* is tempera-484 ture in kelvins, and *R* is the universal gas constant. *Hirth* 485 and Kohlstedt [1996, 2003] showed that for olivine either 486 saturated with hydrogen ("wet") and with small amounts of 487 hydrogen ("dry"), *n* differs little. The addition of hydrogen 488 reduces the effective viscosity largely by reducing *A*, by 489 nearly 2 orders of magnitude. Here, we ignore the high-490 stress, low-temperature flow law suggested by *Evans and* 491 *Goetze* [1979] which shows lower effective values of *B* than 492



**Figure 3.** Average values B(T) in mantle lithosphere as function of temperature at the Moho, bounds of average values B(T) obtained from applying scaling laws for Rayleigh-Taylor instability to conditions in the Altiplano (horizontal lines), and corresponding ranges of Moho temperatures (vertical lines). We integrated values of *B* appropriate for "wet" and "dry" olivine and for eclogite from 1600 K (asthenosphere) to different values of temperature at the Moho. For "wet" olivine (red line) we assumed n = 3,  $H_a = 420$  kJ mol<sup>-1</sup>, and  $A = 1.9 \times 10^3$  MPa<sup>-3</sup> s<sup>-1</sup>, for "dry" olivine (blue line) n = 3.5,  $H_a = 540$  kJ mol<sup>-1</sup>, and  $A = 2.4 \times 10^5$  MPa<sup>-3.5</sup> s<sup>-1</sup> [*Hirth and Kohlstedt*, 1996, 2003], and for eclogite (green line) n = 3.4,  $H_a =$ 480 kJ mol<sup>-1</sup>, and  $A = 2.0 \times 10^3$  MPa<sup>-3.4</sup> s<sup>-1</sup> [*Jin et al.*, 2001].

those given by (6) when temperatures drop below  $\sim 900$  K 493[e.g., Evans and Goetze, 1979; Goetze, 1978]. We recognize 494that (6) cannot be extrapolated reliably to low temperatures, 495a result corroborated by deformation using a diamond anvil 496 [Wenk et al., 2004]. In the temperature range of interest 497 here, however, including the high-stress flow law will alter 498inferred Moho temperatures by amounts much smaller than 499those due to uncertainties in B. 500

[29] We may calculate average values of *B* by integrating 501502B(T) in (6) from the temperature at the Moho to 1600 K, a reasonable temperature for the base of the lithosphere. To 503carry out this integration, we assume that the temperature 504gradient across the mantle lithosphere is constant, and we 505used parameters in flow laws for both "wet" and "dry" 506olivine [Hirth and Kohlstedt, 1996, 2003] and for eclogite 507[Jin et al., 2001] (Figure 3). A comparison of the calcu-508lations based on flow laws with the ranges of values of B509given above based on the scaling laws for Rayleigh-Taylor 510instability, and for n = 3.5, suggest temperatures at the 511Moho of  $\sim$ 814–905 K (540–630°C) for "wet" olivine and 512 $\sim$ 910–1015 K (640–740°C) for "dry" olivine or eclogite 513(Figure 3). Allowance for an uncertainty of a factor of 2 in 514the values of B, estimated either from laboratory measure-515ments or from the scaling laws and geological observations 516widens the range of temperatures by  $\sim 50$  K to  $490-680^{\circ}$ C 517and 590-790°C. The upper of these bounds differs little 518

from the Moho temperature that *Morency and Doin* [2004] 519 inferred to be necessary for removal of mantle lithosphere 520 obeying a plastic constitutive relationship. 521

[30] These ranges of inferred temperatures call for a 522 relatively cool upper mantle (including eclogite) beneath 523 the Altiplano before mantle lithosphere was removed, but if 524 in fact, a thick layer of eclogitized lower crust were 525 removed, such values are not unreasonable for thickened 526 lower crust. For instance, if normal lithosphere with a 527 temperature at the Moho of 490-790°C doubled in thick- 528 ness between  $\sim 40$  and  $\sim 10$  Ma, then at 10 Ma, the 529 temperature structure would have been slowly evolving 530 toward a warmer crust and uppermost mantle, because of 531 the doubled thickness of radiogenic heat production in it. 532 With the long thermal time constant of thickened litho- 533 sphere, however, the temperature near the Moho should 534 have changed little during this 30-Myr span, a result 535 corroborated by numerical calculations of Babevko et al. 536 [2002]. 537

[31] This simple scenario of homogenous thickening 538 overlooks any interaction of the lower lithosphere with the 539 downgoing slab, which clearly could affect the thermal 540 structure of the overlying crust. It also overlooks the 541 possibility that some or all of the mantle lithosphere was 542 removed earlier, for instance, at  $\sim 25$  Ma. Widespread 543 magmatism spanning the width of the present-day plateau 544 seems to have begun near this time [e.g., Allmendinger et 545 al., 1997; James and Sacks, 1999]. On the basis of the 546 westward sweep of magmatic activity, James and Sacks 547 [1999] associated this magmatism with a steepening of 548 Nazca slab subduction from nearly flat to its present-day 549 dip of  $\sim 30^{\circ}$ . Such a change in the dip of the subducting slab 550 beneath the Western Cordillera might have led to some 551 removal of Andean lithosphere, but the occurrence of 552 volcanism implies that the slab was deeper than  $\sim 120$  km. 553 Thus it does not imply that the Andean lithosphere was 554 unusually thin. 555

[32] Most measurements of heat flux from the Altiplano 556 show higher than normal values [Henry and Pollack, 1988; 557 Springer and Förster, 1998; Uyeda and Watanabe, 558 1982], with Henry and Pollack [1988] giving an average 559 of 84 mW/m<sup>2</sup>. Values as large as 84 mW/m<sup>2</sup> would suggest 560 steady state Moho temperatures above 1000°C if relatively 561 typical distributions of radiogenic heat production were 562 assumed [e.g., Springer, 1999], but for a thickened layer of 563 radiogenic heat production and the same surface heat flux, 564 equilibrium Moho temperatures 700°C would also be possi- 565 ble. Because of the transient temperature distribution in the 566 crust, however, the heat flux measurements place only weak 567 constraints on the Moho temperature at  $\sim 10$  Ma. We consider 568 the range of 490-790°C to be cooler than most would expect 569 for a subduction zone, but not unreasonably cool. Babeyko et 570 al. [2002], in fact, suggest that convection within the crust 571 has altered the thermal structure considerably. 572

#### 5. Conclusions

573

[33] Recent analyses of stable isotopes [*Garzione et al.*, 574 2006; *Ghosh et al.*, 2006] in authigenic carbonates in the 575

Altiplano corroborate previous inferences of a post-10 Ma 576rise of the Andean plateau [e.g., Allmendinger et al., 1997; 577 Gregory-Wodzicki, 2000; Gubbels et al., 1993; Isacks, 5781988; Kay et al., 1994; Lamb and Hoke, 1997] and suggest 579that an elevation change of  $3\pm 1$  km occurred in only  $3\pm 1$ 5801 Myr between  $\sim$ 10 and 6.8 Ma. The combination of such a 581large change in mean elevation and the rapidity with which 582it occurred virtually requires removal not only of all of the 583mantle lithosphere but also a layer of dense eclogite beneath 584585the overlying crust [Garzione et al., 2006]. [34] We use scaling laws that relate the time for a 586 perturbation to the thickness of dense layer overlying a less 587 dense layer to grow to infinite depth and the measured time 588of  $3 \pm 1$  Myr to estimate the average viscosity coefficient for 589the mantle lithosphere that underlay the Altiplano before 59010 Ma and then was removed:  $1.4 (\pm 0.7) \times 10^{11}$  Pa s<sup>1/3.5</sup> B < 8.8 (±4.4) × 10<sup>11</sup> Pa s<sup>1/3.5</sup>. The change in elevation of 3 ± 591592

1 km constrains the mass per unit area of the layer that was removed, so that the largest sources of uncertainty in the estimated average viscosity coefficient are the boundary condition at the top of the layer and the approximation of lithosphere as a layer of constant viscosity used to derive the

598 scaling laws.

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[35] A comparison of the range of estimates of viscosity 599 coefficient with those based on calculated averages from 600 temperature-dependent constitutive relations based on lab- 601 oratory measurements yield a range of likely temperatures 602 between  $\sim 500^{\circ}$ C and  $\sim 800^{\circ}$ C at the Moho of the Altiplano 603 at  $\sim 10$  Ma, just before mantle lithosphere was removed. 604 Such values are consistent with those appropriate for 605 doubling the thickness of the lithosphere between  $\sim 40~606$ and  $\sim 10$  Ma. In addition, this range of average values of 607 viscosity coefficient implies average deviatoric stresses 608 within the mantle lithosphere less than  $\sim 50$  MPa. Thus, 609 insofar as these scaling laws for Rayleigh-Taylor instability 610 can be applied to removal of mantle lithosphere, the 611 corresponding material parameters match those deduced 612 from laboratory experiments and show that high strengths 613 of earth-forming minerals offer no obstacle to removal of 614 tectonically thickened mantle lithosphere. 615

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